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# Entropy conserving implicit time integration in a Discontinuous Galerkin solver in entropy variables

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# Abstract

This article presents a fully discrete entropy conserving/stable method based on a Discontinuous Galerkin (DG) discretization in entropy variables coupled with a modified Crank-Nicolson scheme. The entropy conserving time integration is inspired by the work of LeFloch [1], originally developed in the context of a Finite Volume method in conservative variables. This entropy conserving time integrator is here adapted to a DG discretization in entropy variables also demonstrating the fulfilment of entropy conservation regardless of the time step size and the type of elements used (quadrangular or triangular elements, possibly with curved edges). The performance of the implicit method will be demonstrated by computing several inviscid flow problems, *i.e.*, the convection of an isentropic vortex, the double shear layer, the Kelvin-Helmholtz instability, the shedding flow past a triangular wedge, the Sod shock tube, the receding flow and the Taylor-Green vortex.

*Keywords:* Discontinuous Galerkin, generalized Crank-Nicolson, entropy conserving/stable discretizations, entropy variables.

## 1 1. Introduction

<sup>2</sup> This paper presents an entropy conserving method for the numerical solution

<sup>3</sup> of the Euler equations in the context of a high-order Discontinuous Galerkin (DG)

<sup>4</sup> discretization. Hyperbolic systems of partial differential equations (PDEs), such as

<sup>5</sup> the one describing the behaviour of compressible inviscid flows, can admit several

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weak solutions. To properly select the physical relevant one an entropy function 6 can be defined. This function is constant where the solution is smooth and it can 7 only increase across singularities such as shock waves. The idea to embed this 8 physical constraint in the numerical method is not new, see for example the seminal 9 works of Tadmor et al. [2, 3], but has received increasing attention in recent years, 10 in particular when applied in the context of high-order DG methods. This class 11 of numerical methods is very attractive for Computational Fluid Dynamics (CFD) 12 thanks to the ability to get high-accuracy on unstructured and hybrid grids [4] and 13 the favourable dissipation and dispersion proprieties that makes DG well suited to 14 the scale-resolving simulation of turbulent flows, e.g., [5, 6, 7]. 15

DG discretizations with entropy conserving capabilities usually follow two ap-16 proaches: i) the use of Summation by Parts (SBP) operators, split forms of the Eu-17 ler equations and the conservative set of variables, see the works of Gassner et al., 18 e.g., [8, 9]; *ii*) the use of a symmetrized form of the governing equations together 19 with the set of entropy variables. The latter method, employed in this work, does 20 not introduce any limitation at the continuum level but it requires to "exactly" eval-21 uate all the integrals part of the discrete form. This can be numerically achieved 22 by means of *over-integration*, *i.e.*, by computing integrals using quadrature for-23 mulas with a degree of exactness large enough to make negligible the integration 24 errors of non-polynomial functions, see [10]. When numerical methods that use 25 a piecewise discontinuous representations of the solution are considered, properly 26 27 designed numerical flux functions that guarantee entropy conservation/stability must be used, see for example [11, 12, 13]. Most of them were proposed in the 28 context of low-order Finite Volume (FV) entropy conserving schemes. Space-time 20 DG methods can also be considered to devise entropy conserving schemes. How-30 ever, in the authors' opinion, such approach is impractical, as all the time slabs 31 are linked together by a centred temporal state (numerical flux in time), e.g., see 32 Friedrich *et al.* [14]. For this reason, the method of lines is here considered, which 33 implies the use of numerical fluxes and time integration schemes both having en-34 tropy conserving properties. 35

While the development of specifically designed numerical fluxes has been the 36 subject of several works, less attention has been devoted in literature to the develop-37 ment of entropy conserving time integrators. Recent articles from Lozano [15, 16] 38 show that, for a generic entropy function, explicit and fully implicit Runge-Kutta 39 schemes introduce spurious entropy. The papers also indicate the Backward-Euler 40 method as an entropy stable scheme, *i.e.*, the entropy evolution in time is mono-41 tone. Gouasmi et al. [17] also show that for both the BDF2 and the explicit 42 Leap-Frog methods it is difficult to determine a priori the sign of entropy pro-43

duction, while the explicit Forward Euler scheme is entropy unstable. In Colombo *et al.* [10] no clear statement about the entropy production of the linearly implicit
Runge-Kutta schemes of the Rosenbrock type was given. Note that, differently
from [15, 16, 17], in [10] the spatial discretization is based on a DG method with
entropy variables and not on more "standard" FV spatial approximations.

Up to the authors knowledge, the only entropy conserving scheme available in 49 the literature is a modified version of the Crank-Nicolson method. This modified 50 scheme is due to LeFloch et al. [1] and is often refereed as "Generalized Crank-51 *Nicolson*" method. The scheme was originally developed in the context of FV 52 but is considered impractical as it requires numerical quadrature to assemble the 53 modified intermediate state which substitutes the algebraic mean in the residuals 54 vector evaluation of the standard scheme. Gouasmi et al. [17] proposed a compu-55 tationally efficient implementation of the method by using theoretical arguments 56 which are very similar to those used in the development of entropy conserving 57 flux functions. It is worth mentioning that a similar idea was already proposed 58 in Subbareddy and Candler [18] to obtain a fully discrete FV scheme capable of 59 preserving kinetic energy. 60

The extension of these results to DG discretizations is not straightforward but 61 it can be considered of great interest. In fact, an efficient and high-order entropy 62 stable numerical framework that is also essentially dissipation-free in time, is an 63 excellent candidate for explicit LES as it allows a sharp control on the amount of 64 artificial dissipation added by the subgrid-scale model [19]. This paper will report 65 in detail how to implement the entropy conserving time integrator scheme in a DG 66 modal solver, in particular when entropy variables are used. The change of vari-67 ables, in fact, involves a projection error that must be carefully taken into account. 68 Theoretical and numerical proofs will demonstrate that the approach is entropy 69 conserving in the sense that the  $L_2$ -projection of the entropy variables on the con-70 servative ones results in a "global" entropy conservation up to machine precision 71 irrespective of the time step size. The main features of the proposed method are: 72 i) the use of a couple of  $L_2$ -projections between the entropy and the conservative 73 variables and vice versa; *ii*) to retain the time-derivative of the conservative vari-74 ables in the governing equations instead of using the time derivative of the entropy 75 ones. This is done even though the solution is sought in terms of the degrees of 76 freedom of the entropy variables, a fundamental difference from the approach fol-77 lowed by the authors in [10]. As in this work the conservation form of the Euler 78 equations is considered, feature *ii*) guarantees a conservative discretization. 79 In this paper, together with the implementation details of the method, numeri-80

cal results for several test cases will be presented to assess the performance of the

entropy conserving framework regardless of the number and the type of elements,
 the spatial order of accuracy and the time step size. Results will be also presented
 for the entropy stable method, obtained by considering properly designed numer-

<sup>85</sup> ical flux functions, and for "standard" time-integration schemes.

# **2.** The governing equations

The set of equations governing the behaviour of inviscid flows, *i.e.*, the Euler equations, can be written for the compressible case as

$$\begin{aligned} \frac{\partial \rho}{\partial t} &+ \frac{\partial}{\partial x_j} \left( \rho u_j \right) = 0, \\ \frac{\partial}{\partial t} \left( \rho u_i \right) &+ \frac{\partial}{\partial x_j} \left( \rho u_j u_i \right) = -\frac{\partial p}{\partial x_i}, \\ \frac{\partial}{\partial t} \left( \rho E \right) &+ \frac{\partial}{\partial x_j} \left( \rho u_j H \right) = 0, \end{aligned}$$
(1)

where  $\rho$  is the fluid density,  $\mathbf{u} = \{u_1, \dots, u_d\}$  the velocity vector, *E* and *H* the total energy and enthalpy, and  $i, j = 1, \dots, d$ , where *d* is the number of geometrical dimensions. For a perfect gas, the pressure *p* is given by  $p = (\gamma - 1) \rho \left[ E - (u_i u_i)/2 \right]$ , where  $\gamma = c_p/c_v$  is the ratio of gas specific heats, here set to 1.4. In compact form the system (1) can be written as

$$\frac{\partial \mathbf{q}}{\partial t} + \frac{\partial \mathbf{F}_i(\mathbf{q})}{\partial x_i} = \mathbf{0},\tag{2}$$

with implied summation on index *i* and where  $\mathbf{q} = \{\rho, \rho u_i, \rho E\}$  is the vector of the conservative variables and  $\mathbf{F} \in \mathbb{R}^{2+d} \otimes \mathbb{R}^d$  is the convective flux function.

#### 96 2.1. The entropy framework

The concept of entropy conservation relies on the existence of a convex function, the *generalized entropy function* S, and the scalar-valued *entropy flux* functions  $U_i$ , such that the following compatibility conditions holds

$$\frac{\partial S(\mathbf{q})}{\partial \mathbf{q}} \frac{\partial \mathbf{F}_i(\mathbf{q})}{\partial \mathbf{q}} = \frac{\partial \mathcal{U}_i(\mathbf{q})}{\partial \mathbf{q}} \qquad \text{with} \qquad i = 1, \dots, d.$$
(3)

<sup>100</sup> By left-multiplying Eq. (2) by  $\partial S(\mathbf{q})/\partial \mathbf{q}$  we obtain

$$\frac{\partial S(\mathbf{q})}{\partial t} + \frac{\partial U_i(\mathbf{q})}{\partial x_i} = 0, \tag{4}$$

with implied summation on index i. Since entropy needs to be dissipated at shock waves, the above identity is replaced by the following inequality

$$\frac{\partial S(\mathbf{q})}{\partial t} + \frac{\partial \mathcal{U}_i(\mathbf{q})}{\partial x_i} \le 0, \tag{5}$$

that, assuming periodic boundary conditions, can be integrated in space giving

$$\frac{d}{dt} \int_{\mathbb{R}^d} \mathcal{S}(\mathbf{q}) d\mathbf{x} \le 0.$$
(6)

Note that S differs from the common thermodynamic definition of entropy, which increases across shocks.

By assuming that *S* is strictly convex, the mapping  $\mathbf{q} \rightarrow \partial S(\mathbf{q})/\partial \mathbf{q}$  can be regarded as a change of variables from the conservative variables  $\mathbf{q}$  to the entropy variables  $\mathbf{v} = \partial S(\mathbf{q})/\partial \mathbf{q}$ . The system of governing equations (2) is symmetrized when rewritten in terms of entropy variables as

$$\frac{\partial \mathbf{q}(\mathbf{v})}{\partial \mathbf{v}} \frac{\partial \mathbf{v}}{\partial t} + \frac{\partial \mathbf{F}_i(\mathbf{q}(\mathbf{v}))}{\partial \mathbf{v}} \frac{\partial \mathbf{v}}{\partial x_i} = \mathbf{0},\tag{7}$$

where  $\partial \mathbf{q}(\mathbf{v}) / \partial \mathbf{v}$  is positive-definite.

According to Huges *et al.* [20] the only *generalized entropy function-entropy flux pair*, up to a multiplicative constant, which symmetrizes the viscous term in the compressible Navier–Stokes equations is

$$S = -\frac{\rho s}{\gamma - 1}, \quad U_i = S u_i, \quad \text{whith} \quad i = 1, \dots, d,$$
 (8)

where  $s = \ln(p\rho^{-\gamma})$  is the thermodynamic entropy. The corresponding set of entropy variables is

$$\mathbf{v} = \left\{ \frac{\gamma - s}{\gamma - 1} - \frac{\rho}{2p} |\mathbf{u}|^2, \, \frac{\rho u_i}{p}, \, -\frac{\rho}{p} \right\}.$$
(9)

Since  $\partial \mathbf{q}(\mathbf{v})/\partial \mathbf{v}$  and  $\partial \mathbf{F}_i(\mathbf{q}(\mathbf{v}))/\partial \mathbf{v}$  are symmetric, there exists an *entropy potential*,  $\vartheta$ , *entropy flux potential*,  $\psi_i$ , *pair* such that  $\partial \vartheta/\partial \mathbf{v} = \mathbf{q}(\mathbf{v})$  and  $\partial \psi_i/\partial \mathbf{v} = \mathbf{F}_i(\mathbf{q}(\mathbf{v}))$ . It can be easily proved that  $\vartheta = \mathbf{v} \cdot \mathbf{q}(\mathbf{v}) - S(\mathbf{q}(\mathbf{v}))$  and  $\psi_i = \mathbf{v} \cdot \mathbf{F}_i(\mathbf{q}(\mathbf{v})) - U_i(\mathbf{q}(\mathbf{v}))$ , and that they reduce to  $\vartheta = \rho$  and  $\psi_i = \rho u_i$  when the generalized entropy function of Eq. (8) is used.

In this work, following the seminal work of Huges *et al.* [20], the system of governing equations (7) is solved by directly approximating the entropy variables v in the discrete space. To fulfil entropy conservation at the discrete level, the approach assumes that integrals must be computed exactly [21]. However, this statement can be relaxed as shown in [10]. Indeed, numerical results demonstrate that if integrals are approximated by using "accurate enough" quadrature rules, the entropy conservation is verified.

#### **3.** The Generalized Crank-Nicolson scheme in the DG framework

In this work, when "standard" time integration schemes are considered, they are applied to the governing equations written in the following form

$$\frac{\partial \mathbf{q}(\mathbf{v})}{\partial \mathbf{v}} \frac{\partial \mathbf{v}}{\partial t} + \frac{\partial \mathbf{F}_i(\mathbf{q}(\mathbf{v}))}{\partial x_i} = \mathbf{0}.$$
 (10)

By applying the classical Crank-Nicolson (SCN) scheme the system of semi-discrete
 governing equations becomes

$$\frac{\partial \mathbf{q}(\mathbf{v}^{n+1/2})}{\partial \mathbf{v}} \frac{\mathbf{v}^{n+1} - \mathbf{v}^n}{\Delta t} + \frac{\partial \mathbf{F}_i(\mathbf{q}(\mathbf{v}^{n+1/2}))}{\partial x_i} = \mathbf{0},\tag{11}$$

where  $\mathbf{v}^{n+1/2}$  is defined according to the algebraic mean of the working variables, *i.e.*,  $1/2(\mathbf{v}^{n+1} + \mathbf{v}^n)$ . Equation (11) is not in a conservative form due to the Jacobian of the change of variables,  $\partial \mathbf{q}(\mathbf{v}^{n+1/2})/\partial \mathbf{v}$ .

LeFloch *et al.* [1] designed the following entropy conserving time integration scheme

$$\frac{\mathbf{q}^{n+1} - \mathbf{q}^n}{\Delta t} + \frac{\partial \mathbf{F}_i(\mathbf{q}(\mathbf{v}^{n+1/2}))}{\partial x_i} = \mathbf{0},\tag{12}$$

138 where

$$\mathbf{v}^{n+1/2}\left(\mathbf{q}^{n},\mathbf{q}^{n+1}\right) = \int_{0}^{1} \mathbf{v}\left((1-\chi)\mathbf{q}^{n}+\chi\mathbf{q}^{n+1}\right)d\chi.$$
 (13)

The scheme was proposed in the context of FV entropy conserving schemes for
non-linear systems of equations. Following the Tadmor [2] terminology this method
will be referred as "Generalized Crank-Nicolson" (GCN). It is proven that when

GCN is coupled with an entropy conserving flux function, the resulting FV dis-142 cretization is entropy conserving both in space and in time. However, it is not 143 obvious how to extend this result to a high-order DG discretization in entropy 144 variables. In fact, in Eq. (10) the set of working variables is not the conservative 145 but the entropy one and the time derivative term is multiplied by the Jacobian ma-146 trix  $\partial q(\mathbf{v})/\partial \mathbf{v}$ . Moreover, the projection between the entropy variables, v, and the 147 conservative ones,  $\mathbf{q}$  (and vice versa) involves an approximation error. To solve 148 these issues Eq. (7) has been recast in the traditional conservative form as 149

$$\frac{\partial \mathbf{q}(\mathbf{v})}{\partial t} + \frac{\partial \mathbf{F}_i(\mathbf{v})}{\partial x_i} = \mathbf{0},\tag{14}$$

where, with an abuse of notation,  $\mathbf{F}_i(\mathbf{v})$  is used in place of  $\mathbf{F}_i(\mathbf{q}(\mathbf{v}))$ . The time discretization of Eq. (14) gives

$$\frac{\mathbf{q}(\mathbf{v}^{n+1}) - \mathbf{q}(\mathbf{v}^n)}{\Delta t} + \frac{\partial \mathbf{F}_i(\mathbf{v}^{n+1/2})}{\partial x_i} = \mathbf{0}.$$
 (15)

The system of governing equations is now discretized in space by multiplying Eq. (15) by an arbitrary smooth test function  $\mathbf{w} = \{w_1, \dots, w_{d+2}\}$  and integrating by parts, to obtain the weak formulation of each k-th scalar equation

$$\int_{\Omega} w_k \frac{q_k(\mathbf{v}^{n+1}) - q_k(\mathbf{v}^n)}{\Delta t} d\Omega = \int_{\Omega} \frac{\partial w_k}{\partial x_i} F_{k,i}\left(\mathbf{v}^{n+1/2}\right) d\Omega - \int_{\partial\Omega} w_k F_{k,i}\left(\mathbf{v}^{n+1/2}\right) n_i d\sigma, \qquad (16)$$

where  $\Omega \in \mathbb{R}^d$ , with  $d \in \{2,3\}$ , is the domain,  $\partial\Omega$  its boundary and  $\mathbf{n} = \{n_1, \dots, n_d\}$  the unit vector normal to the boundary. In this work, when used as a subscript, the symbol k is an index spanning the range  $1, \dots, 2 + d$ , otherwise it will represent the degree of polynomial functions.

Denoting by  $\mathcal{K}_h = \{K\}$  a mesh of the domain made of elements K, the solution 159 in conservative,  $\mathbf{q}$ , and entropy,  $\mathbf{v}$ , variables, together with the test function  $\mathbf{w}$ , are 160 replaced with finite element approximations,  $\mathbf{q}_h$  and  $\mathbf{v}_h$ , and a discrete test function 161  $\mathbf{w}_h$ , all of them belonging to the discrete polynomial space  $[\mathbb{P}_d^k(\mathcal{K}_h)]^{2+d}$ . Each 162 component  $q_{h,k}$  of the numerical solution can be expressed, in terms of the elements 163 of the global vector **Q** of unknown degrees of freedom, as  $q_{h,k} = \phi_i Q_{k,i}$ , with 164  $j = 1, ..., (N_{dof}) \operatorname{card}(\mathcal{K}_h)$ , where  $N_{dof} = \prod_{i=1,d} (k + i/i)$  and  $\phi_i$  belongs to the 165 set of orthogonal and hierarchical basis functions defined according to [4]. For the 166

sake of notation, the set  $\mathcal{F}_h \stackrel{\text{def}}{=} \mathcal{F}_h^i \cup \mathcal{F}_h^b$  of mesh faces is defined, where  $\mathcal{F}_h^b$  collects the faces lying on the boundary of  $\Omega_h$ . For any  $F \in \mathcal{F}_h^i$  there exist two elements 167 168  $K^+, K^- \in \mathcal{K}_h$  such that  $F \in \partial K^+ \cap \partial K^-$ . For all  $F \in \mathcal{F}_h$ ,  $\mathbf{n}_F$  is the normal unit 169 vector pointing from  $K^+$  to  $K^-$ , for the sake of notation compactness, the subscript 170 F will be dropped in the following. Since a function  $w_h \in \mathbb{P}^k_d(\mathcal{K}_h)$  is double valued 171 over an internal face  $F \in \mathcal{F}_h^i$ , the jump trace operator  $[[w_h]] \stackrel{\text{def}}{=} w_{h|K^+} - w_{h|K^-}$ 172 is defined. This operator acts componentwise when applied to a vector. The DG 173 discretization of the Euler equations results in seeking, for k = 1, ..., 2+d, the 174 elements of **Q** such that 175

$$\sum_{K \in \mathcal{K}_{h}} \int_{K} \phi_{i} \phi_{j} \frac{Q_{k,j}^{n+1} - Q_{k,j}^{n}}{\Delta t} d\Omega = \sum_{K \in \mathcal{K}_{h}} \int_{K} \frac{\partial \phi_{i}}{\partial x_{l}} F_{k,l} \left( \mathbf{v}_{h}^{n+1/2} \right) d\Omega$$
$$- \sum_{F \in \mathcal{F}_{h}} \int_{F} \left[ \left[ \phi_{i} \right] \right] \widehat{F}_{k} \left( \mathbf{v}_{h}^{\pm n+1/2}, \mathbf{n} \right) d\sigma.$$
(17)

To demonstrate the conserving/stability properties of the present scheme we focus firstly on the discrete form of the unsteady term of Eq. (4) resulting from the LHS of (17).  $Q_{k,j}^n$  and  $Q_{k,j}^{n+1}$  are evaluated as the  $L_2$ -projection of the conservative variables (computed according to the entropy variables) on the DG polynomial space

$$Q_{k,i}^{n} = [(\mathbf{M}^{K})^{-1}]_{i,j} \int_{K} \phi_{j} q_{k}(\mathbf{v}_{h}^{n}) d\Omega, \qquad (18)$$

where  $\mathbf{M}_{i,j}^{K} = \int_{K} \phi_{i} \phi_{j} d\Omega$  is the mass matrix of the element *K*. The projected conservative variables  $\hat{q} \in [\mathbb{P}_{d}^{k}(\mathcal{K}_{h})]^{2+d}$  are then defined as

$$\widehat{q}_{h,k}^n = \phi_j Q_{k,j}^n. \tag{19}$$

183 Similarly,

$$V_{k,i}^{n+1/2} = [(\mathbf{M}^K)]_{i,j}^{-1} \int_K \phi_j v_k^{n+1/2} d\Omega,$$
(20)

184 and

$$v_{h,k}^{n+1/2} = \phi_j V_{k,j}^{n+1/2}.$$
(21)

When  $v_{h,k}^{n+1/2} = \phi_i V_{k,i}^{n+1/2}$  are used as the test functions in place of  $\phi_i$  in Eq. (17), for each  $K \in \mathcal{K}_k$  the unsteady term can be re-formulated as

$$\int_{K} \phi_{i} V_{k,i}^{n+1/2} \phi_{j} \frac{Q_{k,j}^{n+1} - Q_{k,j}^{n}}{\Delta t} d\Omega = \left( \int_{K} \phi_{i} \phi_{j} d\Omega \right) V_{k,i}^{n+1/2} \frac{Q_{k,j}^{n+1} - Q_{k,j}^{n}}{\Delta t} =$$

$$= \mathbf{M}_{j,i}^{K} [(\mathbf{M}^{K})^{-1}]_{i,m} \left( \int_{K} \phi_{m} v_{k}^{n+1/2} d\Omega \right) \frac{Q_{k,j}^{n+1} - Q_{k,j}^{n}}{\Delta t} =$$

$$= \delta_{j,m} \left( \int_{K} \phi_{m} v_{k}^{n+1/2} d\Omega \right) \frac{Q_{k,j}^{n+1} - Q_{k,j}^{n}}{\Delta t} = \int_{K} v_{k}^{n+1/2} \phi_{j} \frac{Q_{k,j}^{n+1} - Q_{k,j}^{n}}{\Delta t} d\Omega =$$

$$= \int_{K} v_{k}^{n+1/2} \frac{\hat{q}_{h,k}^{n+1} - \hat{q}_{h,k}^{n}}{\Delta t} d\Omega,$$
(22)

where  $\delta_{i,m}$  is the Kronecker delta function. The LeFloch's [1] identity states that

$$v_k^{n+1/2}\left(\widehat{q}_{h,k}^{n+1} - \widehat{q}_{h,k}^n\right) = \mathcal{S}(\widehat{\mathbf{q}}_h^{n+1}) - \mathcal{S}(\widehat{\mathbf{q}}_h^n),\tag{23}$$

and here it holds true pointwise if  $\mathbf{v}^{n+1/2}$  is defined according to the projection of the entropy variables on the conservative set, *i.e.*,  $\hat{\mathbf{q}}_h^n$  and  $\hat{\mathbf{q}}_h^{n+1}$ . In other words, Eq. (23) is valid when the intermediate state is

$$v_k^{n+1/2}\left(\widehat{\mathbf{q}}_h^n, \widehat{\mathbf{q}}_h^{n+1}\right) = \int_0^1 v_k\left((1-\chi)\,\widehat{\mathbf{q}}_h^n + \chi\,\widehat{\mathbf{q}}_h^{n+1}\right) d\,\chi. \tag{24}$$

<sup>191</sup> For the sake of compactness, in the following,  $v_k^{n+1/2}(\hat{\mathbf{q}}_h^n, \hat{\mathbf{q}}_h^{n+1})$  will be denoted <sup>192</sup> as  $v_k^{n+1/2}$ , accordingly,  $\mathbf{v}^{n+1/2}$  was used in place of  $\mathbf{v}^{n+1/2}(\mathbf{q}^n, \mathbf{q}^{n+1})$ . Equation (24) <sup>193</sup> implies that the last integral of Eq. (22) is equivalent to

$$\int_{K} \frac{S(\hat{\mathbf{q}}_{h}^{n+1}) - S(\hat{\mathbf{q}}_{h}^{n})}{\Delta t} d\Omega,$$
(25)

which means that the PDE for the entropy evolution is implicitly discretized in a conservative form and the fully discrete scheme is entropy conserving. Moreover, since the conservative form of the governing equations is used, the scheme is fully conservative in time, *i.e.*, mass, momentum and energy are conserved at the discrete level. It is worth noting that  $S(\mathbf{q}(\mathbf{v}_h^n))$  is different from  $S(\widehat{\mathbf{q}}_h^n)$ . In fact, the peculiarity of the scheme is that while  $V_{k,j}^{n+1}$  are the DOFs, the scheme is entropy conserving in the sense that the projection on the conservative set is entropy conserving, so the entropy solution is based on the  $Q_{k,j}^{n+1}$  coefficients. However, during the simulations Eqs. (18) and (19) are only used for the evaluation of Eq. (24) since Eq. (22) is solved implicitly. Notice that, the  $L_2$ -projection procedure does no alter the order of accuracy of the method, neither in space nor in time.

From the implementation point of view, Eq. (24) requires numerical integration and to properly select the degree of exactness of the quadrature rules to guarantee entropy conservation, a value difficult to be estimated *a priori*. A viable alternative was proposed by Gouasmi *et al.* [17] who found a close form solution of Eq. (23). The idea behind this method is very similar to that used for the derivation of entropy conserving numerical flux functions. The intermediate state  $\mathbf{v}^{n+1/2}$ of Gouasmi *et al.* is computed as

$$v_{1}^{n+1/2} = \frac{1}{\gamma - 1} \left( \gamma \frac{\overline{\rho}}{\overline{\rho^{ln}}} - \overline{s_{h}} \right) - \overline{u_{i,h}} v_{1+i}^{n+1/2} - \frac{1}{2} \overline{u_{i,h}} u_{i,h}^{n+1/2}, 
 v_{1+i}^{n+1/2} = -\overline{u_{i,h}} v_{2+d}^{n+1/2} 
 v_{2+d}^{n+1/2} = -\frac{\overline{\rho_{h}}}{\overline{p_{h}^{ln}}},$$
(26)

where i = 1, ..., d and where  $p_h = (\gamma - 1)[(\rho E)_h - (\rho u_i)_h (\rho u_i)_h / (2\rho_h)]$ ,  $u_{i,h} = (\rho u_i)_h / \rho_h$  and  $s_h = \ln(p_h \rho_h^{-\gamma})$  are evaluated using the projected solutions  $\hat{\mathbf{q}}_h^n$  and  $\hat{\mathbf{q}}_h^{n+1}$ . The arithmetic and logarithmic means are defined as

$$\overline{a_h} = \frac{a_h^{n+1} + a_h^n}{2}, \quad \overline{a_h^{\ln}} = \frac{a_h^{n+1} - a_h^n}{\ln(a_h^{n+1}) - \ln(a_h^n)}.$$
(27)

Gouasmi *et al.* [17] also proved that the resulting scheme is second-order accurate in time. In the following this method will be denoted as GCNG. Here, thanks to the use of orthonormal shape functions, see [4] for the details, the implementation strongly simplifies reducing the computational cost for the evaluation of the intermediate state  $\mathbf{v}_h^{n+1/2}$ .

Friedrich *et al.* [14] used a similar approach to derive an entropy conservative *temporal state* for their space-time DG algorithm. Their method takes advantage of the SBP property, it uses the conservative variable set and does not require *over-integration*. Their *temporal state* is essentially a central flux function defined in time that links together all the time slabs. The resulting space-time entropy conservative scheme is not a real option since the only viable choice is a pure upwind *temporal state* which decouples the time slabs. However, in this case the scheme is "only" entropy stable.

<sup>228</sup> Concerning the RHS of (17), the linear combination with the coefficients  $V_{k,i}^{n+1/2}$ , <sup>229</sup> where k = 1, ..., 2 + d and  $i = 1, ..., (N_{dof}) \operatorname{card}(\mathcal{K}_h)$ , gives

$$\sum_{K \in \mathcal{K}_h} \int_K \frac{\partial v_k^{n+1/2}}{\partial x_l} F_{k,l}\left(\mathbf{v}_h^{n+1/2}\right) d\Omega - \sum_{F \in \mathcal{F}_h} \int_F \left[ \left[ v_{h,k}^{n+1/2} \right] \right] \widehat{F}_k\left(\mathbf{v}_h^{\pm n+1/2}, \mathbf{n}\right) d\sigma.$$
(28)

By defining the potential flux as  $\psi_l = v_k F_{k,l} - U_l$ , it can be shown that  $\partial v_k / \partial x_l F_{k,l} = \partial U_l / \partial x_l + \partial \psi_l / \partial x_l - v_k \partial F_{k,l} / \partial x_l$ , cf. [21]. Moreover, since the compatibility condition holds, *i.e.*,  $\partial U_l / \partial x_l = v_k \partial F_{k,l} / \partial x_l$ , it is possible to write  $\partial v_k / \partial x_l F_{k,l} = \partial \psi_l / \partial x_l$ . For the sake of notation compactness, in the above equation the dependence on  $\mathbf{v}_h^{n+1/2}$  was omitted, and in the following the discrete potential flux  $\psi_l(\mathbf{v}_h^{n+1/2})$  will be denoted as  $\psi_{h,l}^{n+1/2}$ . By substituting the previous result into Eq. (28) the following relation is obtained

$$\sum_{K \in \mathcal{K}_{h}} \int_{K} \frac{\partial \psi_{h,l}^{n+1/2}}{\partial x_{l}} d\Omega - \sum_{F \in \mathcal{F}_{h}} \int_{F} \left[ \left[ v_{h,k}^{n+1/2} \right] \right] \widehat{F}_{k} \left( \mathbf{v}_{h}^{\pm n+1/2}, \mathbf{n} \right) d\sigma, \qquad (29)$$

<sup>237</sup> which, by using the divergence theorem, becomes

$$\sum_{F \in \mathcal{F}_h} \int_F \left[ \left[ \left[ \left[ \psi_{h,l}^{n+1/2} n_l \right] \right] - \left[ \left[ v_{h,k}^{n+1/2} \right] \right] \widehat{F}_k \left( \mathbf{v}_h^{\pm n+1/2}, \mathbf{n} \right) \right] d\sigma.$$
(30)

As an entropy conserving numerical flux function fulfils the following relation, cf. [21],

$$\left[\left[\boldsymbol{\psi}_{h,l}^{n+1/2}\boldsymbol{n}_{l}\right]\right] - \left[\left[\boldsymbol{\upsilon}_{h,k}^{n+1/2}\right]\right]\hat{F}_{k}\left(\mathbf{v}_{h}^{\pm n+1/2},\mathbf{n}\right) = 0,\tag{31}$$

Eq. (30) is identically zero and the scheme is entropy conserving in space. Similarly, using an entropy stable numerical flux where

$$\left[\left[\boldsymbol{\psi}_{h,l}^{n+1/2}\boldsymbol{n}_{l}\right]\right] - \left[\left[\boldsymbol{v}_{h,k}^{n+1/2}\right]\right]\widehat{F}_{k}\left(\mathbf{v}_{h}^{\pm n+1/2},\mathbf{n}\right) \ge 0,\tag{32}$$

an entropy stable space DG discretization is obtained. Notice that, these results
assume an exact evaluation of all the integrals. This result is not new, see [22], but
it is here demonstrated on a different perspective. Together with Eq. (25) it proves

that the proposed method implicitly discretizes the entropy conservation/inequality
of Eq. (6),

$$\sum_{K \in \mathcal{K}_{h}} \int_{K} \frac{S(\widehat{\mathbf{q}}_{h}^{n+1}) - S(\widehat{\mathbf{q}}_{h}^{n})}{\Delta t} d\Omega = \sum_{F \in \mathcal{F}_{h}} \int_{F} \left[ \left[ \left[ \psi_{h,l}^{n+1/2} n_{l} \right] \right] - \left[ \left[ v_{h,k}^{n+1/2} \right] \right] \widehat{F}_{k} \left( \mathbf{v}_{h}^{\pm n+1/2}, \mathbf{n} \right) \right] d\sigma \ge 0.$$
(33)

Finally, the analogy between Eq. (31) and Eq. (23) highlights the similarity between the entropy conserving flux functions,  $\hat{\mathbf{F}}(\mathbf{v}^{\pm n+1/2}, \mathbf{n})$ , and the intermediate state in time,  $\mathbf{v}^{n+1/2}$ .

#### 250 3.1. Implementation details

Numerical integration of Eq. (17) by means of suitable Gauss quadrature leads
 to the following discrete system of non-linear equations

$$\mathbf{M}\frac{\mathbf{Q}(\mathbf{V}^{n+1}) - \mathbf{Q}(\mathbf{V}^n)}{\Delta t} + \mathbf{R}(\mathbf{V}^{n+1/2}) = \mathbf{0},$$
(34)

where, according to Eq. (20) and Eq. (24) or Eq. (27),  $\mathbf{V}^{n+1/2}$  is a non-linear function of  $\mathbf{V}^{n+1}$  and  $\mathbf{V}^n$ .

Eq. (34) is here solved with a Newton-Kyrolv algorithm. For the sake of clarity, we will use here the notation  $\mathbf{V}^{n+1/2} = \hat{\mathbf{V}}(\mathbf{V}, \mathbf{V}^n)$ , where  $\mathbf{V}$  can be the actual solution vector of the Newton algorithm,  $\mathbf{V}^{(i)}$ , or the solution vector at the end of the time step,  $\mathbf{V}^{n+1}$ . The, possibly inexact, *i*-th Newton step finds  $\Delta \mathbf{V}^{(i)} = (\mathbf{V}^{(i+1)} - \mathbf{V}^{(i)})$ such that

$$\left(\frac{\mathbf{M}}{\Delta t}\frac{\partial \mathbf{Q}(\mathbf{V}^{(i)})}{\partial \mathbf{V}} + \frac{\partial \mathbf{R}}{\partial \widehat{\mathbf{V}}}\frac{\partial \widehat{\mathbf{V}}}{\partial \mathbf{V}}\right)\Delta\mathbf{V}^{(i)} = -\mathbf{M}\frac{\mathbf{Q}(\mathbf{V}^{(i)}) - \mathbf{Q}(\mathbf{V}^{n})}{\Delta t} - \mathbf{R}, \quad (35)$$

where the dependence of **R** and  $\partial \mathbf{R} / \partial \hat{\mathbf{V}}$  on  $\hat{\mathbf{V}}(\mathbf{V}^{(i)}, \mathbf{V}^n)$  was omitted for the sake of notation compactness.

The matrix  $\partial \hat{\mathbf{V}} / \partial \mathbf{V}$  is a block-diagonal matrix which couples all the DOFs of an element and, particularly for the GCN, is not trivial to be derived and implemented. For this reason, this contribution to the implicit operator was approximated as (1/2)**I**, where **I** is the identity matrix. The idea was to simply handle the term as in the standard Cranck-Nicolson scheme. According to our numerical experiments, the impact of this approximation on the non-linear convergence was small, *i.e.*, the number of non-linear steps required by GCN or GCNG to reach a
full, machine precision, convergence was almost the same as for the SCN, which
uses an exact Jacobian matrix. An example will be given in Sec. 4.1.

In all the implicit time integration schemes used in this paper, the Newton algorithm starts with  $\mathbf{V}^{(1)} = \mathbf{V}^n$  and stops when  $||\Delta \mathbf{V}^{(k)}||_{L^2}$  reaches the tolerance  $\varepsilon_{tol}$ . For the solution of the linear systems arising from both implicit and linearly implicit Rosenbrock-type schemes the preconditioned GMRES solvers available in the PETSc library are used [23].

#### **4. Numerical results**

The performance of the entropy conserving Crank-Nicolson scheme named 277 GCNG in Sec. 3, in terms of accuracy of the solution and conservation property, 278 has been assessed and compared to other time integrators, by computing the fol-279 lowing inviscid flow problems: i) the isentropic vortex; ii) the double shear layer; 280 *iii*) the Kelvin-Helmholtz instability; *iv*) the shedding flow past a triangular wedge; 281 v) the Sod shock tube; vi) the receding flow; vii) the Taylor-Green vortex. For the 282 first flow problem, which has an exact analytical solution, the accuracy will be 283 measured in terms of the  $\eta$  error, while the conservation properties, evaluated for 284 several test cases, will be evaluated in terms of the  $\varepsilon$  error, which are defined as 285 follows 286

$$\eta(\circ, \bullet_{ref}) = (\Omega_h)^{-1/2} ||\circ - \bullet_{ref}||_{L^2},$$
(36)

$$\varepsilon(\circ, \bullet_{ref}) = \Omega_h^{-1} \left( \int_{\Omega_h} \circ d\Omega - \int_{\Omega_h} \bullet_{ref} d\Omega \right), \tag{37}$$

<sup>287</sup> where  $\circ$  and  $\bullet_{ref}$  are the numerical and the "reference" solutions, respectively. The <sup>288</sup> reference value is set equal to the  $L_2$ -projection of the initial solution on the DG <sup>289</sup> polynomial space.

The results obtained with the GCN and/or the GCNG schemes will be com-290 pared with the computations from other time integrators. In particular, with the 291 acronym BE we denote the first-order accurate Backward Euler scheme, with SCN 292 the second-order accurate standard Crank-Nicolson scheme, with BDF2 the second-293 order accurate Backward Differentiation Formulae, with ROSXY the linearly-implicit 294 Rosenbrock-type Runge-Kutta schemes, where X is the order of accuracy and Y is 295 the number of stages, with RK35 the third-order accurate Strong Stability Preserv-296 ing explicit Runge-Kutta scheme with 5 stages, with FE the first-order accurate 297

explicit Forward Euler scheme. Details on the implementation and the efficient use of Rosenbrock-type schemes in the DG framework can be found in [24, 25].

For all the simulations using implicit time integration schemes, to avoid any influence of the time integration error, the non-linear solver tolerance  $\varepsilon_{tol}$  has been set to  $10^{-13}$ , whereas the linear solver tolerance has been set to  $10^{-2}$  to speed-up the iterative process. For the Rosenbrock-type schemes the iterative linear solver tolerance was set equal to  $10^{-13}$ .

To obtain a fully discrete entropy conserving/stable scheme, dedicated nu-305 merical flux functions with entropy conserving/stable properties have been im-306 plemented. In particular, the entropy conserving (EC) numerical flux of Ismail 307 and Roe [11], the kinetic energy preserving and entropy-conservative (KEEC) flux 308 of Chandrashekar [12], the entropy stable numerical flux based on the entropy-309 consistent dissipation of Roe (ES), named EC1 in [11], and the entropy stable Go-310 dunov flux computed from the exact solution of the Riemann problem (ERS) [26] 311 have been used. Following the findings of Colombo et al. [10], when simulations 312 are performed with the entropy preserving fluxes, *over-integration* must be con-313 sidered to ensure entropy conservation. However, in the authors' experience, its 314 role becomes less relevant for simulations of simple flow problems that uses the 315 ES and ERS fluxes. In practice, all the computations shown in this paper were 316 performed using *over-integration*, with the exception of the isentropic vortex and 317 the double shear layer (on Cartesian grids) cases. 318

In the next sections, when reported, the elemental CFL number is defined as  $CFL_{K} = (|\mathbf{u}_{a}| + c_{a})\Delta t/\Delta x_{K}$ , where  $c_{a} = \sqrt{\gamma p_{a}/\rho_{a}}$  is the speed of sound and  $\Delta x_{K} = d(V_{K}/S_{K})$  is a measure of the cell size where  $V_{K}$  and  $S_{K}$  are the cell volume (area when d = 2) and surface area (perimeter when d = 2), respectively. The *a* subscript in the CFL definition means that variables are computed from elemental mean values.

#### *4.1. The isentropic vortex convection problem*

The convection of an inviscid isentropic vortex [27, 28, 29] is here considered to assess the accuracy and the entropy conservation properties of the Crank-Nicolson scheme named GCNG in Sec. 3, using the GCN scheme as reference. The initial flow condition is given by

329

$$u_{1} = U - \frac{\alpha}{2\pi} \left( x_{2} - \frac{L}{2} \right) e^{\beta(1-r^{2})},$$
  

$$u_{2} = U + \frac{\alpha}{2\pi} \left( x_{1} - \frac{L}{2} \right) e^{\beta(1-r^{2})},$$
  

$$\Theta = 1 - \frac{\alpha^{2} (\gamma - 1)}{16\beta\gamma\pi^{2}} e^{2\beta(1-r^{2})},$$
  

$$p = 1,$$
(38)

where  $\Theta$  is the temperature and the "free-stream" non dimensional velocity com-330 ponents are equal to  $U = \sqrt{\gamma}$ , corresponding to a Mach number  $M_{\infty} \approx 1.4$ . The 331 symbol r is the distance of a generic point of the computational domain of co-332 ordinates,  $(x_1, x_2)$ , with respect to the vortex center, placed in the middle of the 333 computational domain. The  $\alpha$  and  $\beta$  values are set equal to 5 and 1/2, respectively. 334 The computational domain is  $0 \le x_i \le L$ , with L = 10, discretized with a fine 335 and a coarse uniform mesh made of  $50 \times 50$  and  $25 \times 25$  quadrilateral elements, 336 respectively. Boundary conditions are periodic and simulations are performed up 337 to a final time T, corresponding to one period of vortex revolution. 338

Figure 1 shows a temporal convergence study performed using the fine grid, the 339  $\mathbb{P}^3$  DG polynomial approximation and the EC flux. The largest time step size re-340 ported in the plots corresponds to CFL  $\approx$  3.8. Figure 1(a) shows the performance 341 of the two entropy conserving Crank-Nicolson schemes compared with other im-342 plicit and linearly-implicit temporal integrator showing that all of them achieve 343 their formal order of accuracy. Moreover, for each of the time step size consid-344 ered, the schemes SCN, GCN and GCNG show the same  $\eta$  error values. As all the 345 variables show similar behaviours, the  $\eta$  error for the first variable only is reported 346 from now on. The  $\varepsilon$  error for  $\rho s$ , reported in Fig. 1(b), shows that both the GCN 347 and GCNG schemes reach the machine precision independently from the time step 348 size, while all the other time integrators show convergence rates equal to or larger 340 than the formal ones. The  $\varepsilon$  error is evaluated, for both the entropy conserving 350 time integration schemes, by using the  $L_2$ -projection of the entropy variables on 351 the conservative ones, cf. Sec. 3. Due to the definition of  $\epsilon$ , see Eq. (37), the con-352 vergence rates can be mainly ascribed to the dissipation error. The measured rates 353 are order 3 for BDF2, and order 4 for SCN, while BE, ROS33 and ROS58 show a 354 rate equal to the theoretical one, *i.e.*, 1, 3 and 5, respectively. For the GCN scheme 355 a Gaussian quadrature rule corresponding to a very high degree of exactness (27) 356 was used to compute the intermediate state of Eq (13) in order to guarantee the en-357 tropy conservation independently from the time step size. Although a comparison 358

of the computational effort for the different time integrators is beyond the scope 359 of this section, it is clear that the GCN always requires a larger CPU time than the 360 GCNG. In fact, the GCNG replaces the explicit quadrature by the evaluation of 361 much cheaper algebraic formulae. For example, when considering the small time 362 step size  $\Delta t = T/2560$ , the use of a third-degree accurate quadrature formula 363 allows to reach the zero machine precision value of S ( $\mathbb{P}^3$ , fine grid, EC flux). 364 However, even using this low-degree rule, the GCN is  $\approx 1.7$  times slower than the 365 GCNG. Increasing the degree of exactness to 27, the safe choice suitable for very 366 large  $\Delta t$  sizes, the CPU time becomes  $\approx 5.5$  times larger. 367

When comparing the performance of the GCNG and SCN schemes, a similar 368 number of inexact Newton's iterations, 7 vs. 6, and almost the same CPU time, the 369 GCNG is 5% faster, is measured. These values prove that the the computational 370 cost needed by the GCNG to compute  $V^{n+1/2}$ , corresponding to the 12.6% of the 371 overall CPU time, is counterbalanced by the assembly and inversion of the change-372 of-variables matrix used by the SCN scheme, cf. Eq. (11) and Eq. (15). Although 373 these results may vary considerably, e.g., depending on the  $\Delta t$  size and the degree 374 of the polynomial approximation, the same conclusion can be drawn: the proposed 375 GCNG scheme can be less computationally demanding than a standard method, 376 such as the SCN. 377



Figure 1: Isentropic vortex problem – Time refinement study. Simulations performed on the fine grid using the  $\mathbb{P}^3$  DG approximation, the EC numerical flux and several time integration schemes.

Figure 2 shows a time refinement study for both the coarse and fine grids using the  $\mathbb{P}^3$  approximation. In particular, Fig. 2(a) displays  $\eta$  computed for the first

conservative variable,  $q_1$ , obtained trough the  $L_2$ -projection of the entropy set on 380 the conservative one. For large time step sizes the computations on the two meshes 381 show the same error values and achieve the formal order of accuracy of the scheme. 382 When the time step is reduced the temporal error dominates the spatial one and 383 different plateau values are reached. Notice that, the minimum error for  $q_1$  on the 384 fine mesh is lower than the one for  $v_1$  shown in Fig. 1(a). The error on entropy in 385 Fig. 2(b) is directly computed from the working variables and, as expected, does 386 not verify the entropy preservation, although using a fully conserving scheme, cf. 387 Sec. 3. 388



Figure 2: Isentropic vortex problem – Time refinement study. Simulations performed on the fine and the coarse grid using the  $\mathbb{P}^3$  DG approximation, the EC numerical flux and the GCNG time integration scheme.

Figure 3 reports a time refinement study performed on the coarse grid using 389 the  $\mathbb{P}^3$  DG approximation and the entropy conserving and entropy stable numerical 390 fluxes, EC and ES, respectively. In this plot the results from the GCNG method 391 are compared with those from the ROS33 scheme. The simulations performed 392 with the ES flux do not use *over-integration*. In fact, for this flow problem, it was 393 observed that even using a large number of quadrature points leads to the same 394 results. As expected, in Fig. 3, when the spatial error overwhelms the temporal 395 one, the same plateau value for  $\eta$  is obtained by the different time integrators for 396 a given numerical flux. Although for a large enough time step size the ROS33 397 scheme reaches a significantly lower  $\eta$  error level than the GCNG, the linearly-398 implicit scheme cannot guarantee the entropy conservation for an arbitrary step 390

size but only for very small values, *i.e.*, when the temporal error is negligible with respect to the spatial one. On the contrary, the GCNG scheme reaches both the plateau values for  $\varepsilon$ ,  $\approx 10^{-7}$  for the ES flux and the machine precision for the EC flux, independently from the time step size, thus demonstrating the entropy preserving capability of the time integrator.



Figure 3: Isentropic vortex problem – Time refinement study. Simulations performed on the coarse grid using the  $\mathbb{P}^3$  DG approximation, the EC and ES numerical fluxes and the ROS33 and the GCNG time integration schemes.

Based upon to the findings of this section, among the two generalized CN methods considered in this work, only the GCNG scheme was used for the computation of the other test cases as it combines the entropy-conserving property with an affordable implementation suitable for production runs.

#### 409 4.2. The inviscid double shear layer

The inviscid double shear layer [30, 31] flow problem is used to assess the conservation properties of entropy, kinetic energy and enstrophy. The kinetic energy is defined as  $\kappa = \rho u_i u_i/2$ , and the enstrophy as  $\zeta = \omega^2$ , where  $\omega$  is the vorticity. Long time simulations performed with the GCNG time integration scheme, using different time step sizes and the EC and ES numerical fluxes, have been considered. The initial flow condition is given by

$$u_{1} = \begin{cases} U \tanh \left[ \left( x_{2} - \pi/2 \right) / \delta_{1} \right] & \text{if } x_{2} \leq \pi, \\ U \tanh \left[ \left( 3\pi/2 - x_{2} \right) / \delta_{1} \right] & \text{if } x_{2} > \pi, \end{cases}$$

$$u_{2} = U \delta_{2} \sin(x_{1}), \qquad (39)$$

$$p = 1,$$

$$\rho = 1,$$

where  $U = M_{\infty} \sqrt{\gamma}$ ,  $\delta_1 = \pi/15$  and  $\delta_2 = 0.05$ , and, to obtain a nearly incompress-416 ible flow,  $M_{\infty} = 0.01$ . The computational domain  $0 \le x_i \le 2\pi L$ , with L = 1, 417 has been discretized by: i) a  $8 \times 8$  uniform Cartesian mesh; ii) an anisotropic 418 and not uniform mesh made of 176 triangular elements; *iii*) a  $8 \times 8$  mesh made of 419 quadrangular elements with curved parabolic edges. At all the boundaries periodic 420 conditions are imposed and the simulations are advanced in time up to  $T = 80T_c$ , 421 where  $T_c = L/U$  is the convective time. The  $T_c$  value is ten times larger than the 422 one usually used in the literature. 423

Figure 4 shows the evolution in time of the errors on entropy, kinetic energy 424 and enstrophy for the Cartesian grid using the GCNG time integration scheme 425 coupled with the EC numerical flux and several time step sizes for the  $\mathbb{P}^{3,4}$  DG 426 approximations. To quantify the errors related to enstrophy and kinetic energy 427 conservation, their relative percentage variation with respect to the reference value, 428 *i.e.*,  $\varepsilon_r(\zeta, \zeta_{ref})$ % and  $\varepsilon_r(\kappa, \kappa_{ref})$ %, is monitored. Figure 4(a) shows an error on 429 entropy close to machine precision for any time step value, thus confirming the 430 entropy conserving capability of the GCNG scheme. Plots 4(b) and 4(c) reveal 431 that, for long time simulations, the even polynomial degree approximations show 432 a significant dissipation for both  $\kappa$  and  $\zeta$  when compared to the odd case, this 433 odd/even effect was already observed and commented by the authors in [10]. This 434 behaviour contributes to a worse stability of  $\mathbb{P}^4$  than  $\mathbb{P}^3$  computations, as testified 435 by the divergent simulations for  $\Delta t = T/200$  and T/400 (not shown in the plots). 436 Note that, the simulations performed with  $\Delta t = T/800$ , which is the larger time 437 step size for which both the odd and even DG approximations lead to a convergent 438 solution, corresponds to CFL  $\approx 25$ . 439

Furthermore, Fig. 4(b) shows that the relative percentage errors on kinetic energy for the  $\mathbb{P}^3$  solutions are almost constant for the different time step sizes considered. Differently, Fig. 4(c) shows that, for the same DG approximation  $\mathbb{P}^3$ , the relative percentage error on enstrophy is greater influenced by the time step size, with error levels that are in the range [2%, 65%], except for the 160% value obtained for the largest  $\Delta t = T/200$  size.



Figure 4: Double shear layers problem – Time refinement study. Simulations performed on the Cartesian grid using the  $\mathbb{P}^{3,4}$  DG approximations, the EC numerical flux and the GCNG time integration scheme with different  $\Delta t$  values.

The same refinement study was performed by using the ES flux and results are summarized in Fig. 5. For the Cartesian grids only, entropy-stable simulations were performed without *over-integration* but verifying that a greater number of quadrature points led to the same results. As expected, cf. [10], computations are not affected by an odd/even effect and the k = 4 approximation shows, especially for entropy and kinetic energy, better conservation properties than the k = 3 case. Differently from the EC case, the use of the ES flux allowed to successfully perform the computation for the  $T/\Delta t = 200$  and  $T/\Delta t = 400$  values.

When using the ES flux combined with the GCNG time integrator, the conservation properties for a given DG approximation are roughly the same for all the time step sizes considered, thus further confirming the good properties of the GCNG. The pressure contours shown in Fig. 6 highlights the different level of accuracy achieved for the  $\mathbb{P}^4$  DG approximation by using two different time steps sizes, *i.e.*,  $T/\Delta t = 200$  and  $T/\Delta t = 6400$ .



Figure 5: Double shear layers problem – Time refinement study. Simulations performed on the Cartesian grid using the  $\mathbb{P}^{3,4}$  DG approximations, the ES numerical flux and the GCNG time integration scheme with different  $\Delta t$  values.



Figure 6: Double shear layers problem – Pressure contours at time  $T = 80T_c$  computed on the Cartesian grid using the  $\mathbb{P}^4$  DG approximation, the ES flux and the GCNG time integration scheme with different  $\Delta t$  values.

To demonstrate the geometrical flexibility of the proposed method, Fig. 7 and 460 Fig. 8 show the results for the EC and the ES fluxes, respectively, using meshes 461 made of triangular elements and quadrilateral elements with curved edges. The 462 grids used, superimposed to the pressure contours, are shown in Fig. 9. The sim-463 ulation that uses triangular elements is performed with CFL  $\approx$  20, while, for the 464 grid with curved edges, this value is  $\approx 17$ . The numerical experiments essentially 465 confirm the findings obtained for the Cartesian mesh case. It can be stated that 466 the method preserves its conservation properties also on unstructured meshes and 467 without the need of any special treatment. 468



Figure 7: Double shear layers problem – Simulations performed on the meshes made of triangular elements (Triangle) and quadrilateral elements with curved edges (Curve), using the  $\mathbb{P}^{3,4}$  DG approximations, the EC numerical flux and the GCNG time integration scheme with  $T/\Delta t = 3200$ .



Figure 8: Double shear layers problem – Simulations performed on the meshes made of triangular elements (Triangle) and quadrilateral elements with curved edges (Curve), using the  $\mathbb{P}^{3,4}$  DG approximations, the ES numerical flux and the GCNG time integration scheme with  $T/\Delta t = 3200$ .



(a) Triangular elements

(b) Curved quadrangular elements

Figure 9: Double shear layers problem – Pressure contours at time  $T = 80T_c$ . Simulations performed on the meshes made of triangular elements and quadrilateral elements with curved edges using the  $\mathbb{P}^4$  DG approximation, the ES flux and the GCNG time integration scheme with  $T/\Delta t = 3200$ .

#### 469 4.3. The Kelvin-Helmholtz instability problem

This section deals with the two-dimensional Kelvin-Helmholtz instability (KHI) 470 problem parametrized in [32] as a function of the Atwood number  $A = (\rho_2 - \rho_2)$ 471  $\rho_1/(\rho_2 + \rho_1)$ . As in the work of Chan et al. [32] an entropy stable DGSEM method 472 (seventh degree polynomial approximation) was observed to begin to crash for 473  $A \approx 0.3$ , to demonstrate the robustness of the present method a value slightly 474 larger was used in the following computations, *i.e.*, A = 1/3. For the sake of com-475 parison with the reference paper, a similar set-up was considered using the entropy 476 stable ERS flux together with a  $32 \times 32$  Cartesian uniform grid and polynomial ap-477 proximations up to  $\mathbb{P}^6$ . 478

KHI-type flow problems are known to be very sensitive to initial conditions
as well as to the numerical resolution and possible small perturbations. In fact,
the flow evolution is characterized by the generation of small structures becoming
smaller and smaller when increasing the discretization accuracy.

It is worth noting that this test case can be considered as a strongly compressible version of the double shear layer problem with a non-uniform density initialization and a much higher velocity magnitude (maximum Mach number equal to 0.6 for t = 0). Although the initial condition can be recast in a form similar to Eq. (39), the domain size,  $-L \le x_i \le L$  with L = 1 and the same definitions reported in [32] are here used

$$u_{1} = U\left(f(x_{2}) - \frac{1}{2}\right),$$

$$u_{2} = \frac{U}{10}\sin(2\pi x_{1}),$$

$$p = 1,$$

$$\rho = \rho_{1} + f(x_{2})\left(\rho_{2} - \rho_{1}\right),$$
(40)

where  $f(x_2) = 1/2 [\tanh(15x_2 + 7.5) - \tanh(15x_2 - 7.5)], U = \sqrt{p/\rho_1} \text{ and } \rho_1 = \frac{1}{2}$ 490 1.

Computations have been integrated in time up to T = 14L/U, a value larger 491 than the final time used in [32]. As the dynamics of the problem is quite fast, the 492 mesh is uniform and the Mach number is high, the computational efficiency of im-493 plicit methods was expected not to be very high. Nevertheless, for the k = 6 case, 494 the use of GCNG allows to march in time with a CFL number 7.5 times larger than 495 the stability limit, *i.e.* 0.2, of the explicit RK35 scheme, see [33], here considered 496 as a reference solution. Figure 10(a) reports the  $\varepsilon(\rho s, \rho s_{ref})$  value as a function 497 of time. For each time integration scheme solutions are shown up to the largest 498 admissible  $CFL_{max}$  value. For the present numerical set-up and among all the con-499 sidered time integrators, *i.e.*, FE, ROS22, ROS33, CNS, GCNG, the ROS33 is 500 the most stable one with a  $CFL_{max} = 2$ . For this value, the same scheme also 501 shows the largest production of entropy on the whole time interval. Concerning 502 the GCNG, the method outperforms both the SCN (CFL<sub>max</sub> = 1) and the ROS22 503 (CFL<sub>max</sub> = 1.25) schemes and results, with CFL<sub>max</sub> = 1.5, in a  $\varepsilon(\rho s, \rho s_{ref})$  profile 504 very close to the reference value, the RK35 with  $CFL_{max} = 0.2$ , or to the result ob-505 tained with a smaller time step size (CFL=0.5). Moreover, the SCN and ROS22 506 schemes proved to be entropy unstable, see the detail in Fig. 10(b). Note that, even 507 if the ROS33 scheme seams here stable, we are not able to guarantee its entropy 508 stability. For example, even the ROS58 scheme proves to be entropy unstable, for 509 large time steps, in Figure 17 of [10]. According to the theory, the stability limit 510 of the Forward Euler is very low and the method exhibits a nonphysical large drop 511 of entropy for t > 8, see Fig. 10(a). These results highlight how an entropy sta-512 ble spatial discretization alone may not guarantee an overall fully discrete entropy 513 stable method. In fact, entropy production in space can be overwhelmed by the 514 entropy destruction in time. 515



Figure 10: Kelvin-Helmholtz instability problem – Comparison of the  $\varepsilon(\rho s, \rho s_{ref})$  evolution obtained with different time integrators and different CFL values. Simulations performed using the  $\mathbb{P}^6$  DG approximation and the ERS numerical flux.



Figure 11: Kelvin-Helmholtz instability problem – Comparison of the  $\varepsilon_r(\kappa, \kappa_{ref})$ % evolution obtained with different time integrators and different CFL values. Simulations performed using the  $\mathbb{P}^6$  DG approximation and the ERS numerical flux.



Figure 12: Kelvin-Helmholtz instability problem – Density contours at the final time T = 14L/U. Simulations performed using the  $\mathbb{P}^6$  DG approximation, the ERS numerical flux and different time integration schemes.

As our objective is to develop a numerical method well suited for the under-516 resolved simulation of multiscale phenomena, e.g., in a LES context, it is of partic-517 ular interest to investigate the kinetic energy conservation properties of the scheme [19]. 518 As expected, the evolution in time of kinetic energy follows a trend similar to en-519 tropy, see Fig. 11, where the  $\varepsilon(\kappa, \kappa_{ref})$ % profile for the GCNG is closer to the 520 reference curves than the ROS33, which significantly dissipates energy, especially 521 for large CFL values. As regards the FE, the nonphysical raise of kinetic energy, 522 shown in Fig. 11, dramatically spoils the solution accuracy, and corresponds to a 523 drop in entropy, as observed in Fig. 10. 524

Notice that, although the theoretical order of accuracy of ROS33 is larger than 525 GCNG, *i.e.*, 3 vs. 2, the density contours shown in Figs. 12(a) and 12(b) reveal that 526 the GCNG delivers a solution containing smaller scales than the ROS33. Obvi-527 ously, this behavior does not mean that the GCNG solution is in absolute more ac-528 curate, see as a reference the solution obtained with the RK35 scheme in Fig. 12(c). 520 However, it highlights the enhanced conservation proprieties of the GCNG, with 530 respects to other schemes, even when using an entropy stable spatial discretization. 531 Finally, Fig. 13 proves that when using the GCNG (CFL = 1) the entropy and 532 kinetic energy time history of lower polynomial degree disctretizations is close to 533 the k = 6 case, although entropy starts to be produced earlier. This behavior is due 534 to the discontinituies of the solution at the mesh faces, which become larger when 535 the numerical accuracy is lowered, as they are the only source of entropy when the 536 ERS flux is coupled with GCNG. 537



Figure 13: Kelvin-Helmholtz instability problem – Time evolution of  $\varepsilon(\rho s, \rho s_{ref})$  and  $\varepsilon_r(\kappa, \kappa_{ref})$ %. Simulations performed using the  $\mathbb{P}^{2\to 6}$  DG approximations, the ERS numerical flux and, if not otherwise indicated, the GCNG time integration scheme.

### 538 4.4. The shedding flow past a triangular wedge

In this section, a simple test case where an inviscid unsteady flow interacts with boundary conditions is presented. The scope of this flow problem is to testify that no particular numerical treatment is needed at the boundary when entropy variables are used. Inspired by [34, 35], the inviscid flow around a triangular wedge is considered for the free-stream condition  $M_{\infty} = 0.2$ . The triangular body immediately generates vortices at sharp corners, even under an inviscid condition, resulting in a vortex shedding behind the wedge.

A natural approach to impose boundary conditions in a DG framework is their weak enforcement, [5], where properly defined states are used, directly or together with the internal states, to compute the numerical fluxes at the boundary faces. These boundary states must be defined according to the condition type and be consistent with the physical flux. In this flow problem inlet/outlet characteristicbased conditions are imposed at left/right boundary, symmetry conditions on the top and bottom boundaries, and the wedge surface is treated as a slip wall.

The computational domain consists of an equilateral triangle with side L = 1, placed on the centerline of a rectangular box of size  $26L \times 10L$  at a distance 6Lfrom the inlet (left) boundary. The simulations were performed using the ERS flux with the  $\mathbb{P}^4$  approximation on an unstructured mesh made of 5 407 triangular elements with linear edges. The solution was advanced in time using the GCNG

scheme with a time step size equal to  $\Delta t = 10^{-2}T_c$  with  $T_c = L/U$ , where U 558 is the free-stream velocity magnitude. During iterations, this time step size 559 corresponds to roughly a CFL value of  $\approx$  2.3. Figure 14 shows the density contours 560 for two different time levels,  $T = 30T_c$  and  $T = 50T_c$ . A density variation of 561 roughly the 30% is observed for this compressible flow problem. For the  $T = 30T_c$ 562 snapshot the mesh has been also superimposed on half of the domain for the sake 563 of completeness. Once generated, the eddies are convected downstream, slightly 564 diffused by the quite coarse mesh, and finally interact with the outlet boundary 565 with only a moderate and expected distortion of the flow structure. 566





Figure 14: Inviscid flow past a wedge – Density contours,  $T = 30T_c$  (top) and  $T = 50T_c$  (bottom). Simulation performed on a grid made of 5 407 triangular elements using the  $\mathbb{P}^4$  DG approximation, the ERS numerical flux and the GCNG time integration scheme.

## <sup>567</sup> 4.5. The receding flow problem

In this section a flow problem where rarefaction is generated by two inviscid flows receding one from each other is considered. This test case has been extensively studied by Liou [36, 37] and used by Gouasmi et al. in their presentation of an entropy conserving time integration scheme [17]. The initial condition is given by two constant states separated by a discontinuity

$$(p, \rho, u) = \begin{cases} (2, 1, -0.4) & \text{if } -0.5 \leq x \leq 0, \\ (2, 1, 0.4) & \text{if } 0 < x \leq 0.5, \end{cases}$$
(41)

note that these values correspond to the ones used in [17]. This test case exhibits 573 a non-physical temperature rise (overheating) with a spurious entropy generation 574 at the origin that cannot be fixed by simply refining the spatial discretization. Liou 575 proposed to cure this problem by replacing the energy conservation equation with 576 a transport equation for the specific entropy [37]. Unfortunately, this approach 577 does not guarantee for the total energy conservation. Gouasmi et al. demonstrated 578 that this spurious entropy rise is also observed in a fully (both in space and in time) 579 entropy conserving numerical scheme [17]. 580

The solutions, here computed up to T = 0.18, have been integrated in time with 581 different values for the time-step size  $\Delta t = \{10^{-3}, 5 \cdot 10^{-4}, 10^{-4}\}$  and comparing 582 the results of the GCNG with the second-order Rosenbrock-type scheme of Iannelli 583 and Baker (ROS22) [38, 24]. Taking advantage of the symmetry condition set at 584 x = 0 the computational domain x = [0, 0.5] is discretized with 50 elements and 585 Dirichlet conditions applied at the right boundary. The temperature profile, with 586 a detail of the symmetry region, is shown for the fully entropy conserving scheme 587 (GCNG and KEEC) and different values of the time step size,  $\Delta t = \{10^{-3}, 5 \cdot \}$ 588  $10^{-4}$ ,  $10^{-4}$ , in Fig. 15. It is worth mentioning that in the authors' experience the 589 EC and KEEC flux functions deliver very similar results on several flow cases [10]. 590 As expected, due to the essentially non-dissipative nature of the scheme, a 591 solution affected by non-physical oscillations is predicted regardless of the time 592 step size, cf. [10]. 593



Figure 15: Receding flow problem – Temperature profiles with detail of the symmetry region (right) at T = 0.18. Simulations performed on the grid composed by 50 elements using the  $\mathbb{P}^6$  DG approximation, the KEEC numerical flux and the GCNG time integration scheme with different values of the time step size ( $\Delta t = \{10^{-3}, 5 \cdot 10^{-4}, 10^{-4}\}$ ).

Figure 16(a) shows a detail of the oscillating entropy profile for the  $\Delta t = 10^{-4}$ value and the  $\mathbb{P}^6$  approximation. Fluctuations strongly reduce when an entropy stable spatial discretization (ERS) and/or the non-entropy-conserving linearly-implicit ROS22 scheme are used.



Figure 16: Receding flow problem – Detail of the specific entropy profiles for the receding flow problem at T = 0.18. Simulations performed on the grid composed by 50 elements using the  $\mathbb{P}^6$  DG approximation, the KEEC and ERS numerical fluxes and the GCNG and ROS22 time integration schemes with  $\Delta t = 10^{-4}$ ).

Temperature profiles are shown in Fig. 17 for the  $\mathbb{P}^4$  approximation and the different values of the time step size when using the ERS flux function together with the GCNG or the ROS22 time integrators. Oscillations are mainly concentrated at the symmetry plane, at the foot of the expansion and reduce when reducing the time step size.



Figure 17: Receding flow problem – Temperature profiles at T = 0.18. Simulations performed on the grid composed by 50 elements using the  $\mathbb{P}^4$  DG approximation, the ERS numerical flux and the GCNG and ROS22 time integration schemes with different values of the time step size  $(\Delta t = \{10^{-3}, 5 \cdot 10^{-4}, 10^{-4}\}).$ 

<sup>603</sup> Focusing on the use of the GCNG scheme and the ERS flux, when increasing <sup>604</sup> the accuracy of the spatial discretization from  $\mathbb{P}^4$  to  $\mathbb{P}^6$ , fluctuations reduces as <sup>605</sup> shown Figure 18. Although pressure is well resolved (Fig. 18(b)), overheating <sup>606</sup> develops with a slightly under-estimated density value (Fig. 18(c)) and a spurious



entropy generation (Fig. 18(d)) at symmetry. This corresponds to the behaviour observed by Gouasmi et al. in [17].

Figure 18: Receding flow problem – Temperature profile for the receding flow problem at T = 0.18 with detail of the symmetry region for pressure, density and entropy. Simulations performed on the grid composed by 50 elements using the  $\mathbb{P}^6$  DG approximation, the ERS numerical flux and the GCNG time integration scheme with  $\Delta t = 10^{-4}$ ).

#### 609 4.6. The Sod shock tube problem

The Sod Shock tube is a Riemann problem for the Euler equations with an initial condition defined as

$$(p, \rho, u) = \begin{cases} (1, 1, 0) & \text{if } -0.5 \le x \le 0, \\ (0.1, 0.125, 0) & \text{if } 0 < x \le 0.5. \end{cases}$$
(42)

The solution, here computed up to  $T = 0.2/\sqrt{\gamma}$ , is made of a left-moving rarefaction wave, a right-moving contact discontinuity and a right-moving shock wave. The computational domain is spatially discretized with 100 elements and symmetry conditions are applied at boundaries. Computations have been performed with the ERS flux, using the  $\mathbb{P}^4$  approximation and integrating the solution in time with 2 000 time steps and the GCNG or the ROS22 [38, 24] schemes.

In high-order spatial discretizations flow discontinuities give rise to spurious 618 oscillations that cause stability issues and a loss of accuracy. To cure this prob-619 lem several strategies to stabilize the solution have been proposed in the literature, 620 e.g., [39, 40, 41, 42]. As in this work the focus is on time integration, to avoid any 621 possible interaction of a discontinuity control algorithm with the solution evolu-622 tion, we opted not to use any shock-capturing approach during computations. This 623 choice leads to spurious oscillations at discontinuities that also propagate to re-624 gions where the solution is almost constant, as shown by the density, temperature 625 and velocity profiles in Fig. 19. 626

The results from the ROS22 and the GCNG schemes (same order of accuracy) show an overall similar behaviour and demonstrate that the present method can be considered robust even in presence of flow discontinuities and when an entropy conserving time integration scheme is employed.



Figure 19: Sod shock tube – Density, temperature and velocity profiles at  $T = 0.2/\sqrt{\gamma}$ . Simulations performed on the grid composed by 100 elements using the  $\mathbb{P}^4$  DG approximation, the ERS numerical flux and the GCNG and ROS22 time integration schemes.

## 631 4.7. The inviscid Taylor-Green vortex

The three-dimensional weakly compressible inviscid Taylor-Green vortex [43] is here considered as a representative problem to evaluate the performance of the GCNG scheme for the scale-resolving simulation of turbulent flows. The initial dimensionless condition is given by

$$u_{1} = U \sin\left(\frac{x_{1}}{L}\right) \cos\left(\frac{x_{2}}{L}\right) \cos\left(\frac{x_{3}}{L}\right),$$

$$u_{2} = -U \cos\left(\frac{x_{1}}{L}\right) \sin\left(\frac{x_{2}}{L}\right) \cos\left(\frac{x_{3}}{L}\right),$$

$$u_{3} = 0,$$

$$p = 1 + \frac{U^{2}}{16} \left[\cos\left(\frac{2x_{1}}{L}\right) + \cos\left(\frac{2x_{2}}{L}\right)\right] \left[\cos\left(\frac{2x_{3}}{L}\right) + 2\right],$$

$$\rho = 1,$$
(43)

where  $U = M_{\infty} \sqrt{\gamma}$ , L = 1 and the "free-stream" Mach number is  $M_{\infty} = 0.1$ . 636 The flow problem is solved on the periodic cube  $0 \le x_i \le 2\pi L$  using two 637 Cartesian grids made of  $8^3$  (coarse) and  $32^3$  (fine) elements, respectively. The 638 simulations have been integrated in time up to  $T = 20T_c$  with CFL  $\approx 1.4$ , where 639  $T_c = L/U$  is the convective time. The inviscid nature of this test case makes 640 it very interesting because an infinite range of scales develop and the numerical 641 solution is under-resolved by definition. Figures 20(a) and 20(b) show the time 642 history of  $\varepsilon$  ( $\rho s$ ,  $\rho s_{ref}$ ) computed with the fine and the coarse grid, respectively, 643 and several DG approximations. These plots distinctly show that a "true" entropy 644 preserving simulation is possible only if the two entropy conserving schemes, EC 645

in space and GCNG in time, are used together. It is worth noting that the computations performed by coupling the ROS33 scheme with the EC or the ERS numerical fluxes deliver a final value of  $\varepsilon(\rho s, \rho s_{ref})$  of the same order of magnitude (10<sup>-4</sup>). This result suggests that the evolution of  $\rho s$  is significantly affected by the time integration scheme. However, surprisingly, the ERS results seem almost insensitive to the choice of the time integrator.

These findings demonstrate the difficulty in *a priori* identifying the trend of entropy when a not provable entropy conserving scheme is used, here the ROS33. In fact, in the authors' experience, for a schemes like the ROS33, the entropy production, and possible destruction, depends on the specific features of the numerical solution.

Figures 20(c) and 20(d) report the time history of the relative percentage variation of kinetic energy and confirm that this quantity is much better conserved when the fully entropy conserving scheme uses odd DG approximations ( $\mathbb{P}^{3,5}$ ) rather than even ones ( $\mathbb{P}^{4,6}$ ), cf. Sec. 4.2. Moreover, although the ROS33 is formally one order more accurate than the GCNG, it is observed that  $\kappa$  is better preserved by the last scheme. In fact, as already noted, the conservation of entropy has a positive



influence in the preservation of kinetic energy ( $\kappa$  should be perfectly conserved for  $M_{\infty} \rightarrow 0$ ).

Figure 20: Taylor-Green vortex problem – Time evolution of  $\varepsilon$  ( $\rho s$ ,  $\rho s_{ref}$ ) and  $\varepsilon_r$  ( $\kappa$ ,  $\kappa_{ref}$ ) %. Simulations performed on the fine grid using the  $\mathbb{P}^{3,4}$  DG approximations (left column) and on the coarse grid using the  $\mathbb{P}^{5,6}$  DG approximations (right column), the EC and ERS numerical fluxes and the GCNG and ROS33 time integration schemes.

#### 665 Conclusion

In this article a fully discrete entropy conserving/stable numerical method for the solution of the Euler equations has been presented. The method uses a DG spatial discretization and a modified Crank-Nicolson time integration scheme. The entropy conserving time integrator, originally proposed in the context of FV schemes,
was adapted to a DG discretization in entropy variables theoretically demonstrating
the fulfilment of entropy conservation regardless of the time step size. The conservation and stability properties have been numerically corroborated by computing
several unsteady compressible flow problems, also considering different types of,
possibly curved, mesh cells.

Future work will be devoted to the solution of the Navier–Stokes equations using the present entropy stable numerical framework with the purpose of addressing scale–resolving simulations of turbulent flows, *e.g.*, DNS, LES.

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